Oscillatory flow in a tube of time-dependent curvature. Part 1. Perturbation to flow in a stationary curved tube

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Motivated by the study of blood flow in the coronary arteries, this paper examines the flow of an incompressible Newtonian fluid in a tube of time-dependent curvature. The flow is driven by an oscillatory pressure gradient with the same dimensionless frequency, $\alpha$, as the curvature variation. The dimensionless governing parameters of the flow are $\alpha$, the curvature ratio $\delta_0$, a secondary streaming Reynolds number $R_s$ and a parameter $R_t$ representing the time-dependence of curvature. We consider the parameter regime $\delta_0 \ll R_t \ll 1$ ($R_s$ and $\alpha$ remain $O(1)$ initially) in which the effect of introducing time-dependent curvature is to perturb the flow driven by an oscillatory pressure gradient in a fixed curved tube. Flows driven by low- and high-frequency pressure gradients are then considered. At low frequency ($\delta_0 \ll R_t \ll \alpha \ll 1$) the flow is determined by using a sequence of power series expansions ($R_s = O(1)$). At high frequency ($\delta_0 \ll R_t \ll 1/\alpha^2 \ll 1$) the solution is obtained using matched asymptotic expansions for the region near the wall (Stokes layer) and the region away from the wall in the interior of the pipe. The behaviour of the flow in the interior is then determined at both small and intermediate values of $R_s$. For both the low and high frequency cases, we find the principal corrections introduced by the time-varying curvature to the primary and secondary flows, and hence to the wall shear stress. The physiological application to flow in the coronary arteries is discussed.

1. Introduction

The main application of the study of unsteady viscous flows in tubes of time-dependent curvature is to blood flow in the coronary arteries. By far the most common arterial disease is atherosclerosis, and in the coronary arteries alone it is the cause of one third of all deaths in Western society. It has been found that atherosclerotic plaques are distributed at preferential sites on the artery walls, such as at the outer wall of an arterial bifurcation, the inside of curves, and regions where the arterial cross-section undergoes an expansion. The sites at which early atherosclerotic plaques are found, and their subsequent development, are now widely believed to be correlated with the distribution of mean arterial wall shear stress and with its pulsatility (Ku et al. 1985; Giddens, Zarins & Glagov 1993; Friedman 1993). It has been observed that plaques tend to develop in regions of low mean wall shear stress and regions where the shear stress changes direction in the course of the cardiac cycle (Caro, FitzGerald & Schroter 1971). Such regions are sites of increased residence time for circulating vasoactive and toxic materials. Thus one mechanism for atherogenesis...
is that the endothelium at these sites is exposed to increased contact with atherogenic agents compared to sites where the wall shear stress is elevated or unidirectional.

We consider the extramyocardial coronary arteries that lie on the surface of the beating heart. As the heart beats they experience both a rotation and a translation. Moreover, their overall geometry, represented by the curvature and torsion of the vessel centreline, as well as their diameter and length, changes with time. In this paper we concentrate on variations in the vessel curvature. The flow of blood in the arteries is pulsatile and as a step towards understanding the properties of unsteady flow in a time-dependent curved tube, we shall examine flows driven by oscillatory pressure gradients.

Considerable theoretical attention has been given to fully developed unsteady flow in fixed uniform curved tubes, of small curvature, driven by a time-dependent pressure gradient. The pioneering analytical study by Lyne (1971) considered a sinusoidally oscillating pressure gradient, of period $T$, with zero mean. At large values of $\alpha (= a(\nu T)^{-1/2})$, where $a$ is the tube radius, Lyne developed a solution by the use of two matched asymptotic expansions. The flow consists of an inviscid core surrounded by a viscous Stokes boundary layer and the Stokes layer thickness, $(\nu T)^{1/2}$, is small compared with the tube radius $a$. Within the Stokes layer a secondary velocity is generated by the nonlinear ‘centrifugal’ force terms in the equations. These secondary motions have a non-zero mean which does not fall to zero at the edge of the Stokes layer but instead drives a (two-vortex) steady secondary streaming in the core. The details of this streaming are dependent upon a secondary streaming Reynolds number $R_s$, but for all values of $R_s$ the direction of the streaming is across the centreline of the tube from the outside of the bend to the inside, the opposite sense to that for steady flow in a curved tube (Dean 1927, 1928).

Mullin & Greated (1980a,b) also analysed oscillatory developing and fully developed flow in a curved tube. They found an analytical solution to the governing equations when $\alpha$ is assumed to be small, corresponding to the case of a low-frequency driving pressure gradient. The leading-order axial and secondary flow is quasi-steady Dean flow.

Subsequent workers have superimposed a mean onto the oscillatory driving pressure gradient in order to investigate how the two types of secondary motion interact (Smith 1975; Blennerhassett 1976). Blennerhassett also found non-uniqueness at certain parameter values. Numerical and experimental analysis of periodic flow in a curved tube has also been carried out by Lin & Tarbell (1980), Rabadi, Simon & Chow (1980); Hamakisotes & Berger (1990) and Zalosh & Nelson (1973). Quantitative studies of pulsatile entry flow in a curved pipe using flow visualization and hot-film anemometry were reported by Chandran, Yearwood & Wieting (1979) as well as Chandran & Yearwood (1981) and later using laser Doppler anemometry by Talbot & Gong (1983).

Non-fully developed steady and unsteady flow has been examined, both in the forms of entry flow with a flat profile (Singh 1974; Yao & Berger 1975), and of flow entering from a straight pipe with a parabolic profile (Smith 1976).

For more detailed reviews and an explanation of further curved tube flows see Pedley (1980) and Berger, Talbot & Yao (1983).

The problem of flow in a tube with time-dependent curvature was first addressed by Lynch, Waters & Pedley (1996). In that paper, the governing equations are derived from the Navier–Stokes and continuity equations using a coordinate system fixed in the frame of reference of the tube. The tube is assumed to be circular in cross-section with a uniform radius that is independent of time. The centreline of the tube is
Oscillatory flow in a tube of time-dependent curvature. Part 1

assumed to be inextensible and points on the tube wall which at one instant lie in a plane perpendicular to the centreline at a point \( S_0 \) remain in that plane as \( S_0 \) moves and the plane rotates. The centreline is also assumed to lie in a constant plane (so the torsion is zero). The blood itself is modelled as an incompressible homogeneous viscous Newtonian fluid. The flows driven by a steady pressure gradient in (i) a tube of uniform but time-dependent curvature; and (ii) a sinuous tube, representing a small oscillation from a straight pipe, were analysed by asymptotic methods.

Here we consider fully developed flow driven by an oscillatory pressure gradient in a tube of uniform time-dependent curvature, as a model for those extramyocardial coronary arteries whose curvature is single signed and varies relatively slowly with longitudinal position. The frequency of the pressure gradient oscillation is taken to be the same as that of the curvature, corresponding to the haemodynamic application; both low and high frequencies will be considered. The mean curvature, equal to the ratio of the pipe radius to the mean radius of curvature of the pipe, is \( \delta_0 \). In §2.1 we state the dimensionless governing equations. In §2.2 we derive two secondary streaming Reynolds numbers, \( R_s \) and \( R_s^1 \), which govern the secondary flow behaviour. Although these Reynolds numbers are applicable at both small and large values of \( \alpha \), their derivation is most clearly achieved by assuming \( \alpha \) to be large. From \( R_s \) and \( R_s^1 \) we define a further parameter \( \mathcal{R}_t \) which, along with \( \delta_0 \), \( R_s \) and \( \alpha \), is an independent governing parameter of the flow. In order that analytical progress can be made, the many governing parameters are taken to be large or small. Thus there exist numerous different cases for which asymptotic solutions to the governing equations may be sought. In §2.3 we consider one particular parameter regime in which the secondary flow consists of a small perturbation to that in a fixed curved tube. (See Waters (1996) for a parameter regime in which the secondary motions are driven by the oscillatory rotation and are almost the same as if the pipe were straight.) In §3 we determine the leading-order governing equations for the chosen parameter regime. In §4 we consider the flow at low frequency (\( R_s \) remains \( O(1) \)). In §5 we examine the high frequency case and determine the behaviour of the flow at both small and intermediate values of \( R_s \). We are particularly interested in the secondary flow, especially the non-zero mean part (steady streaming), and its influence on the axial flow. In both §§4 and 5 the effect of the time-varying curvature upon the secondary streaming and, consequently, on the axial flow is determined. The magnitude and distribution of the resulting wall shear stress is then found. Finally, in §6 we briefly discuss the application of this work to flow in the coronary arteries.

2. Governing equations and parameters

2.1. Full equations

The governing equations for the pressure-driven flow in a tube of time-dependent curvature are derived in the paper by Lynch et al. (1996). Here we give a brief outline of their derivation.

The equations are determined from first principles using the Navier–Stokes and continuity equations for an incompressible viscous fluid:

\[ \nabla \cdot \mathbf{u}' = 0, \quad (2.1a) \]

\[ \frac{\partial \mathbf{u}'}{\partial t} + (\mathbf{u}' \cdot \nabla) \mathbf{u}' = -\frac{1}{\rho} \nabla p + \nu \nabla^2 \mathbf{u}' \quad (2.1b) \]

These equations have been written with respect to an inertial frame of reference.
Figure 1. The curvilinear coordinate system.

\((\hat{x}, \hat{y}, \hat{z})\); \((e_x, e_y, e_z)\) denote the Cartesian coordinate directions, \(\hat{u}'\) is the dimensional velocity vector in this frame, \(\hat{p}\) is the dimensional pressure and \(\hat{t}\) is the dimensional time; \(\rho\) denotes the density and \(\nu\) the kinematic viscosity of the fluid. We are considering the flow in a tube that is both rotating and translating. A non-inertial, curvilinear coordinate system, \((\hat{r}, \theta, \hat{s})\), is chosen that moves and deforms with the pipe: \(\hat{s}\) is the dimensional distance along the centreline from a specified origin, \(O\), and \((\hat{r}, \theta)\) are polar coordinates in the plane of cross-section of the tube (figure 1); \((e_r, e_\theta, e_s)\) denote the curvilinear coordinate directions. The tube wall is taken to be at \(\hat{r} = a\). We consider a change in coordinate system from the inertial Cartesian coordinate system \((\hat{x}, \hat{y}, \hat{z})\), where the centreline moves in the \((\hat{x}, \hat{y})\)-plane, to the non-inertial system \((\hat{r}, \theta, \hat{s})\).

We postulate that one cross-section, at \(\hat{s} = 0\), is fixed with its centre at the origin and the tangent to the centreline locally in the \(e_y\)-direction. The equations are written in terms of the velocity relative to that of the tube wall, so that the no-slip boundary condition for viscous flow at the tube wall can be applied simply; \(\hat{u} = (\hat{u}, \hat{v}, \hat{w})\) are the velocity components in the \((e_r, e_\theta, e_s)\) directions and \(\hat{R}\) is the radius of curvature. The curvature ratio, \(\delta\), is given by \(\delta = a/\hat{R}\). (Here 'hats' denote dimensional quantities.) The variables are non-dimensionalized as follows:

\[
\begin{align*}
(\hat{u}, \hat{v}, \hat{w}) &\rightarrow U_0(u, v, w), \\
(\hat{r}, \hat{s}) &\rightarrow a(r, s), \\
\hat{R} &\rightarrow aR, \\
\hat{p} &\rightarrow \rho U_0^2 p, \\
\hat{t} &\rightarrow Tt,
\end{align*}
\]

where \(U_0\) is a suitable velocity scale and \(T\) is the period of the oscillations. The non-dimensional governing equations are then:

continuity

\[
u_r + \frac{1}{r} u_v + \frac{1}{h} w_s + \frac{\delta \cos \theta}{h} u - \frac{\delta \sin \theta}{h} v + \frac{\nu^2 \delta r}{Re h} \cos \theta = 0,
\]
Oscillatory flow in a tube of time-dependent curvature. Part 1

\[ \frac{\alpha^2}{Re} u_t + uu_r + \frac{1}{r} v u_\theta + \frac{1}{h} w u_s - \frac{1}{r} v^2 
+ \frac{\alpha^4}{Re^2} \cos \theta \left( -\frac{\tilde{\delta}}{\delta^2} + \frac{2\delta^2}{\delta^3} \right) \left( 1 - \cos \delta \right) 
- \frac{\alpha^2}{Re} s^2 \delta \cos \theta \left( 2w + \frac{\alpha^2 s^2 h}{Re} \right) 
- \frac{\delta \cos \theta}{h} w^2 
= -p_r - \frac{1}{Re rh} \left\{ \frac{\partial}{\partial \theta} \left[ \frac{h}{r} (v + rv_r - u_\theta) \right] 
- \frac{\partial}{\partial s} \left[ \frac{1}{h} r (\delta \cos \theta w - hw_\delta + u_\epsilon) \right] \right\} 
- \frac{2\alpha^2 \delta \cos \theta}{Re^2 h}, \tag{2.3b} \]

\[ \frac{\alpha^2}{Re} v_t + vv_r + \frac{1}{r} v v_\theta + \frac{1}{h} w v_s + \frac{1}{r} w 
- \frac{\alpha^4}{Re^2} \sin \theta \left( -\frac{\tilde{\delta}}{\delta^2} + \frac{2\delta^2}{\delta^3} \right) \left( 1 - \cos \delta \right) 
+ \frac{\alpha^2}{Re} s^2 \delta \sin \theta \left( 2w + \frac{\alpha^2 s^2 h}{Re} \right) 
+ \frac{\delta \sin \theta}{h} w^2 
= \frac{1}{r} p_r + \frac{1}{Re h} \left\{ \frac{\partial}{\partial s} \left[ \frac{h}{r} (v + rv_r - u_\theta) \right] 
- \frac{\partial}{\partial r} \left[ \frac{1}{rh} \left( -\delta r \cos \theta w + hw_\delta - rv_s \right) \right] \right\} 
+ \frac{2\alpha^2 \delta \sin \theta}{Re^2 h}, \tag{2.3c} \]

\[ \frac{\alpha^2}{Re} w_t + ww_r + \frac{1}{r} v w_\theta + \frac{1}{h} w^2 + \frac{1}{r} w 
- \frac{\alpha^4}{Re^2} \sin \theta \left( -\frac{\tilde{\delta}}{\delta^2} + \frac{2\delta^2}{\delta^3} \right) \sin \delta 
+ \frac{2\alpha^2}{Re} s^2 \delta \cos \theta u 
- 2\frac{\alpha^2}{Re} s^2 \delta \sin \theta v 
- \frac{2\alpha^2}{Re^2} s^2 \delta h 
+ \frac{\delta \cos \theta}{Re} w v 
+ \frac{\delta \sin \theta}{h} v w 
- \frac{\alpha^2}{Re} \frac{2\delta \cos \theta}{h} w 
= \frac{1}{h} p_r + \frac{1}{Re r} \left\{ \frac{\partial}{\partial r} \left[ \frac{1}{rh} (hw_\delta - \delta r \sin \theta w - rv_s) \right] 
- \frac{\partial}{\partial r} \left[ \frac{1}{h} (hw_\delta - \delta \cos \theta w + u_\epsilon) \right] \right\}. \tag{2.3d} \]

Here

\[ h = [1 + \delta(t) r \cos \theta] \tag{2.4} \]

and is 1/R times the distance from the centre of curvature of the centreline (at a particular s) to the projection on the plane containing the centreline of a general point (r, \( \theta \), s). The boundary conditions are that

\[ u = (u, v, w) = 0 \quad \text{on} \quad r = 1. \tag{2.5} \]

\( z \) is the dimensionless frequency, known as Womersley’s parameter, and \( Re \) is the Reynolds number; they are defined by

\[ z = \left( \frac{a^2}{\nu \lambda} \right)^{1/2} \quad \text{and} \quad Re = \frac{U_0 a}{\nu}. \tag{2.6} \]
In the analysis that follows, it is convenient to follow Lyne (1971) and take the velocity scale $U_0$ equal to $\nu/a$ so that formally $Re = 1$. Physically, $\alpha^2$ is the ratio of the frequency of the oscillations to the inverse time scale for viscous diffusion of momentum across the boundary layer. Thus for small values of $\alpha$ the flow is essentially quasi-steady and is viscous throughout. However, for large values of $\alpha$ the rate of viscous diffusion is small compared with the frequency and the effects of viscous diffusion are confined to thin layers on the boundary.

In equation (2.3b) the term $x^2 \cos \theta (-\ddot{\delta}/\delta^2 + 2\dot{\delta}^2/\delta^3)(1 - \cos s\dot{\delta})$ is the first new term due to time-dependence of curvature. It results from the acceleration of the frame of reference. The next term, $-x^2 s\dot{\delta} \cos \theta 2w$ is a ‘Coriolis’ force associated with the rotation of the frame of reference. The term $-x^2 s\dot{\delta} \cos \theta x^2 (s\dot{\delta}h/\delta)$ is a combination of a centrifugal-force-type term resulting from the rotation of the reference frame and a term due to the acceleration of the frame of reference. The final term in equation (2.3b) is a viscous term associated with the deformation of the frame of reference.

The origin of the new terms is similar for equation (2.3c). In equation (2.3d), the terms $2x^2 s\dot{\delta} \cos \theta u$ and $-2x^2 s\dot{\delta} \sin \theta v$ are ‘Coriolis’ force terms, $(-2x^4 s\dot{\delta}^2/\delta^2$ results from the reference frame acceleration and $(x^2 s\dot{\delta}h)/\delta$ derives from acceleration terms and a centrifugal force term. The additional ‘new’ inertia term is the one linear in $w$ at the end of the third line of (2.3d); this cannot be attributed to any classical ‘forces’ seen in the systems translating and rotating with uniform acceleration and angular velocity, and represents spatial variations in these quantities. All the other terms in equations (2.3b–d) arise also in flow in a fixed curved pipe (see Pedley 1980, chapter 4).

We shall consider the curvature to be

$$\delta = \delta_0(1 + \epsilon \sin t),$$

(2.7)

so that the variation of curvature with time is assumed to be sinusoidal. Here $\delta_0$ is the mean curvature and $\epsilon$ represents the amplitude of its oscillations. We shall take both $\delta_0$ and $\epsilon$ to be small, so that the curvature is small and varies only by a small amount. Note that $\dot{\delta} = \delta_0 \epsilon \cos t$ and $\ddot{\delta} = -\delta_0 \epsilon \sin t$. Now $\delta = 1/R$ and so $\dot{\delta} = -\ddot{R}/R^2$. Thus $\cos t > 0$ corresponds to $\dot{R} < 0$, i.e. to the radius of curvature decreasing with time (negative rotation). Similarly, $\cos t < 0$ corresponds to the radius of curvature increasing with time (positive rotation). The dimensionless axial pressure gradient is given by

$$-p_s = x^2 W_r \cos (t + \Phi),$$

(2.8)

where $W_r$ is the dimensionless velocity amplitude that would be driven by such a pressure gradient in a straight pipe containing an inviscid fluid and $\Phi$ represents the phase difference between the oscillatory pressure gradient and the time-dependent curvature. Note that because the chosen velocity scale is $v/a$, $W_r$ is a Reynolds number for the primary axial flow.

One important feature of curved tube flow is the secondary motion that arises in the cross-sectional plane. This secondary motion itself influences the distribution of axial velocity and results in a complicated interaction of the two. We shall determine the secondary flows that arise due to both the mean and the time-dependent curvature components. We are particularly interested in the non-zero mean part of the secondary flow and its influence on the axial flow and in the magnitude and distribution of the resulting wall shear stress perturbations. The axial and azimuthal components of
Oscillatory flow in a tube of time-dependent curvature. Part 1

Dimensionless wall shear stress are defined as follows:

\[ \text{Axial wall shear stress} = -w_r|_{r=1}, \quad (2.9a) \]

\[ \text{Azimuthal wall shear stress} = -v_r|_{r=1}. \quad (2.9b) \]

2.2. Derivation of the secondary streaming Reynolds numbers when \( \alpha \gg 1 \)

The secondary flow behaviour, in the cross-sectional plane of the tube, is determined by the secondary streaming Reynolds numbers, \( R_s \) and \( R_s_1 \). These are derived in the same way as a conventional Reynolds number except that the velocity scales used are the characteristic secondary velocity scales. The secondary velocity that arises due to the mean curvature of the tube leads to one velocity scale, while the secondary velocity generated by the time-dependence of curvature leads to another. These scales may be found by inspection of equations (2.3a–d) and are the same for both the low and high frequency cases. However, the interpretation of the secondary streaming Reynolds numbers as the controlling parameters for the secondary flow behaviour is most clearly seen by considering the large-\( \alpha \) case, as detailed below. The quantities obtained are equally applicable to the small-\( \alpha \) case.

Lyne (1971) examined flow driven by a high-frequency pressure gradient in a fixed, weakly curved pipe. Here we extend Lyne's analysis to the case of a tube with time-dependent curvature. As in the fixed curved tube case, at large values of \( \alpha \), we expect the flow to consist of an inviscid core surrounded by a viscous Stokes boundary layer, of dimensionless thickness \( O(1/\alpha) \), on the boundary at \( r = 1 \). In the core of the pipe, a balance exists between the centrifugal forces that arise due to the curvature of the pipe, the new terms introduced by the time-dependent curvature and the pressure gradient in that plane. In the Stokes layer the value of \( w \), the axial velocity, must be zero at the wall (no-slip condition). Thus in the boundary layer there will no longer be a balance between the pressure gradient and the centrifugal and Coriolis force terms since the pressure gradient remains essentially unchanged while the latter terms, which are proportional to \( w \), have diminished. Thus oscillatory secondary motions, primarily consisting of a \( \theta \)-component of velocity, \( v \), are generated by the interaction of the local inertia term and the viscous forces with the centrifugal and Coriolis forces respectively. The solution for \( v \) will contain an oscillation of double the fundamental frequency, and a mean flow. These secondary flows must satisfy the no-slip condition at the boundary. However, it is impossible also to impose the condition that these flows tend to zero at the edge of the Stokes layer. Instead, the secondary motions 'drag' the fluid in the core of the tube, thus generating secondary motions there. The details of the core flow will depend on the value of the secondary streaming Reynolds numbers, \( R_s \) and \( R_s_1 \) (see below).

We take the curvature to be small and thus consider a small-parameter expansion in powers of \( \delta_0 \):

\[
\begin{align*}
    u &= u_0 + (2\delta_0)u_1 + \cdots, \\
    v &= v_0 + (2\delta_0)v_1 + \cdots, \\
    w &= w_0 + (2\delta_0)w_1 + \cdots, \\
    p &= p_0 + (2\delta_0)p_1 + \cdots.
\end{align*}
\]

If the viscous terms in equations (2.3a)–(2.3d) are dropped, they have a solution,
satisfying \( u = 0 \) at \( r = 1 \), in which the velocity and pressure distributions are

\[
\begin{align*}
    u &= (2\delta_0) \left[ -\frac{1}{16} x^2 e \cos \theta (1 - r^2) \cos t \right] + O((2\delta_0)^2), \\
    v &= (2\delta_0) \left[ \frac{1}{16} x^2 e \sin \theta (1 - \frac{1}{2} r^2) \cos t \right] + O((2\delta_0)^2), \\
    w &= W_p \sin (t + \Phi) + (2\delta_0) \left[ -2^4 e r \cos \theta s \cos t + O(x^2) \times s \text{-indep. terms} \right] + O((2\delta_0)^2), \\
    p &= (2\delta_0) \left[ x^2 e W_p r \cos \theta s \cos t \sin (t + \Phi) - 2^4 e r \cos \theta \cos t \right] - \frac{1}{16} \alpha^4 e \cos \theta s^2 \sin t + \frac{1}{4} \alpha^2 e r \cos \theta \sin^2 (t + \Phi) \right. \\& \\
    &\left. - \frac{1}{16} \alpha^4 e (r - \frac{1}{4} r^3) \cos \theta \sin t \right] + O((2\delta_0)^2). 
\end{align*}
\]

(2.11)

This can be considered as the primary core flow. In a fixed curved pipe there is no secondary flow in the plane of cross-section as there is a balance between the centrifugal force terms and the pressure gradient in that plane. Here the introduction of time-dependent curvature results in non-zero radial and azimuthal velocities, \( u \) and \( v \). These are necessary in order that the continuity equation, (2.3a), is satisfied at \( O(2\delta_0) \). The fluid pressure adjusts in order to balance the \( s \)-dependent Coriolis forces that arise due to the rotation of the frame of reference, the \( s^2 \)-dependent terms that arise due to the translation of the frame of reference and the viscous terms associated with the deformation of the frame of reference (see equations (2.3b,c)). There is also an adjustment to balance the local inertial terms in (2.3b,c), involving the new radial and circumferential velocities. Note that the secondary velocities and the pressure distribution are \( O(2\delta_0) \) compared with the leading-order axial velocity, and thus are much smaller.

We are interested in determining the magnitude of the secondary velocities that are generated in the boundary layer, since from these the secondary streaming Reynolds numbers can be derived.

For \( \alpha \gg 1 \), we introduce the boundary-layer coordinate

\[ \eta = \frac{x}{\sqrt{2}} (1 - r). \]  

(2.12)

Solutions are sought of the form:

\[
\begin{align*}
    w &= W_0 + (2\delta_0) W_1 + (2\delta_0)^2 W_2 + \cdots, \\
    u &= U_0 + (2\delta_0) U_1 + (2\delta_0)^2 U_2 + \cdots, \\
    v &= V_0 + (2\delta_0) V_1 + (2\delta_0)^2 V_2 + \cdots, \\
    p &= P_0 + (2\delta_0) P_1 + (2\delta_0)^2 P_2 + \cdots;
\end{align*}
\]

(2.13)

\( u, v \) and \( w \) must satisfy the no-slip boundary condition at \( \eta = 0 \), and must match the core flow as \( \eta \to \infty \).

Substitution of (2.12) and (2.13) into the governing equations (2.3a–d) gives the boundary layer equations at successive orders of \( (2\delta_0) \).

The equations at \( O((2\delta_0)^0) \) are just the boundary layer equations for flow in a straight tube driven by an oscillatory axial pressure gradient. The solution to these equations that matches with the core and satisfies the no-slip condition at the boundary is the Stokes layer flow:

\[
\begin{align*}
    U_0 &= 0, \quad V_0 = 0, \quad W_0 = W_p \sin (t + \Phi) - e^{-\eta} \sin (t + \Phi - \eta), \quad P_0 = 0.
\end{align*}
\]

(2.14)

Thus at leading order there are no components of secondary velocity and the axial
oscillatory flow in a tube of time-dependent curvature. Part 1

velocity adjusts to satisfy the no-slip condition. New terms are introduced into the governing equations by the time-dependence of curvature only at $O(2\delta_0)$:

continuity

\[-(z/\sqrt{2})U_{1\theta} + \frac{1}{(1 - (\sqrt{2}/z)\eta)} V_{1\theta} + \frac{1}{(1 - (\sqrt{2}/z)\eta)} U_1 + W_{1s} + \frac{1}{2}(1 + \epsilon \sin t) \cos \theta U_0 - \frac{1}{2}(1 + \epsilon \sin t) \sin \theta V_0 + \frac{1}{2} x^2 \epsilon (1 - (\sqrt{2}/z)\eta) \cos \theta \cos t = 0, \quad (2.15a)\]

and similarly for $e_\theta$,

\[e_s\]

\[x^2 W_{1\theta} - (z/\sqrt{2})(U_0 W_{1\theta} + U_1 W_{0\theta}) + \frac{1}{(1 - (\sqrt{2}/z)\eta)} (V_0 W_{1\theta} + V_1 W_{0\theta}) + W_0 U_{1\theta} + W_1 U_{0\theta} - \frac{1}{2}(1 + \epsilon \sin t) (1 - (\sqrt{2}/z)\eta) \cos \theta W_0 U_0 - \frac{2}{(1 - (\sqrt{2}/z)\eta)} V_0 V_1 + \frac{1}{2} x^2 \epsilon \cos \theta s \sin t - x^2 \epsilon \cos \theta s \cos t W_0 - \frac{1}{2}(1 + \epsilon \sin t) \cos \theta W_0^2 - (z/\sqrt{2}) \frac{\partial P_1}{\partial \eta} - \frac{1}{(1 - (\sqrt{2}/z)\eta)} \frac{\partial}{\partial \theta} \left[ \frac{1}{(1 - (\sqrt{2}/z)\eta)} (V_1 - (z/\sqrt{2}) (1 - (\sqrt{2}/z)\eta) V_{1\theta} - U_{1\theta}) \right] - \frac{\partial}{\partial s} [1 - (\sqrt{2}/z)\eta] (z/\sqrt{2}) (W_{1\theta} + U_{1\theta}) \right] - x^2 \epsilon \cos \theta \cos t, \quad (2.15b)\]

and similarly for $e_\theta$.

An expansion in powers of $(1/z)$ is proposed:

\[
\begin{align*}
W_1 &= x^2 (W_{10} + (1/z) W_{11} + \cdots), \\
U_1 &= x^2 (U_{10} + (1/z) U_{11} + \cdots), \\
V_1 &= x^2 (V_{10} + (1/z) V_{11} + \cdots), \\
P_1 &= x^4 (P_{10} + (1/z) P_{11} + \cdots).
\end{align*}
\]  

(Here we have assumed that $z$ is large, but still $\delta_0 \ll 1/z^2 \ll 1$.)
At leading order equation (2.15a) gives
\[ -\frac{1}{\sqrt{2}} U_{10\eta} = 0 \]  
which, together with the boundary condition that \( U_{10} = 0 \) on \( \eta = 0 \), implies that \( U_{10} \equiv 0 \).

Equation (2.15b) gives that \( P_{10\eta} = 0 \) and so, matching with the pressure in the core given by (2.11), we have
\[ P_{10} = -\frac{1}{8} \epsilon \cos \theta \sin t - \frac{1}{4} \epsilon \cos \theta s^2 \sin t. \]  
(2.18)

The leading-order \( e_\theta \) equation simplifies to
\[ V_{10t} - \frac{1}{2} V_{10\eta\eta} = -\frac{1}{8} \epsilon \sin \theta \sin t. \]  
(2.19)

The solution for \( V_{10} \), that satisfies no-slip and the matching condition given by (2.11), is
\[ V_{10} = \frac{1}{4} \sqrt{2} \epsilon \sin \theta \cos t - \frac{1}{8} \epsilon \cos \theta \cos t, \]  
(2.20)

By considering the axial equation at leading order we find that
\[ W_{10} = -\epsilon \cos \theta s [\cos t - e^{-\eta} \cos (t - \eta)]. \]  
(2.21)

these both represent Stokes layers driven by the \( O(2\delta_0) \) primary core flow.

Similarly, the \( e_r \) and \( e_\theta \) equations, at \( O(\alpha^4) \) and \( O(\alpha^3) \) respectively, give that
\[ P_{11} = \frac{1}{4} \sqrt{2} \epsilon \sin \theta s^2 \sin t, \]  
(2.22)

and
\[ V_{11} = -\frac{1}{8} \sqrt{2} \epsilon \sin \theta \cos t - \frac{1}{16} \sqrt{2} \epsilon \eta \cos \theta \cos (t - \eta). \]  
(2.23)

\( V_{11} \), like \( V_{10} \), is driven by the primary core flow. The continuity equation, (2.15a), at \( O(\alpha^2) \) gives
\[ U_{11} = -\frac{1}{8} \sqrt{2} \epsilon \sin \theta \cos t + \frac{\epsilon}{4} \cos \theta [\sin (t - \eta + \frac{1}{4} \pi) - e^{-\eta} \sin (t - \eta + \frac{1}{4} \pi)]. \]  
(2.24)

Note that \( U_{11} \) does not match the core velocity as \( \eta \to \infty \) (a displacement effect, present in all boundary layers).

Equation (2.15b) at \( O(\alpha^3) \) gives
\[ \frac{1}{\sqrt{2}} P_{12\eta} = \frac{1}{8} \sqrt{2} \epsilon \sin \theta \sin t - \frac{1}{32} \sqrt{2} \epsilon \cos \theta \sin t + \frac{\epsilon}{16} \sqrt{2} \epsilon \cos \theta \cos t, \]  
(2.25)

and so
\[ P_{12} = \frac{1}{8} \epsilon \eta^2 \cos \theta \sin t + \frac{\epsilon}{8} \eta \cos \theta (\cos t - \sin t) - \epsilon \cos \theta \cos t \]
\[ + \epsilon W_p \cos \theta s \sin (t + \Phi) \cos t. \]  
(2.26)

The \( e_\theta \) equation at \( O(\alpha^3) \) is
\[ V_{12r} - \frac{1}{2} V_{12\eta} + \epsilon \sin \theta s \cos t W_0 = \epsilon P_{12r} - \sqrt{2} \eta P_{11\theta} - 2 \eta^2 P_{10\theta} + \text{viscous terms}. \]  
(2.27)

The terms on the right-hand side of (2.27) will lead to a contribution to \( V_{12} \) that is driven by the primary core flow. However, the term \( \epsilon \sin \theta s \cos t W_0 \) will lead to a contribution to \( V_{12} \) that is \( O(\epsilon W_p) \). This contribution to \( V_{12} \) is the first Stokes layer secondary velocity that is not driven by the primary core flow. It is generated by the Coriolis terms that arise because the pipe under consideration has time-dependent curvature.
$P_{13}$ is determined from the radial equation (2.15b) at $O(x^2)$ and at $O(x)$ the equation for $V_{13}$, obtained from the $e_0$ equation is

$$V_{13} - \frac{1}{2} V_{13\eta\eta} + \epsilon \sin \theta \cos t W_{01} = -P_{130} - \sqrt{2} \eta P_{120} - 2 \eta^2 P_{110} - 2 \sqrt{2} \eta^3 P_{100} + \text{viscous terms.} \quad (2.28)$$

Again, the terms on the right-hand side of equation (2.28) will lead to a contribution to $V_{13}$ that is driven by the primary core flow. Moreover, since $W_{01} = 0$, the term on the left-hand side of the equation does not yield any secondary velocities.

Equation (2.15b) at $O(x)$ gives $P_{14}$ and the $e_\theta$ equation at $O(1)$ gives

$$V_{14} - \frac{1}{2} V_{14\eta\eta} + \frac{(1 + \epsilon \sin t) \sin \theta}{2} W_0^2 = -P_{140} - \sqrt{2} \eta P_{130} - 2 \eta^2 P_{120} - 2 \sqrt{2} \eta^3 P_{110} - 4 \eta^4 P_{100} + \text{viscous terms.} \quad (2.29)$$

As before, the terms on the right-hand side lead to a primary core flow driven component of $V_{14}$. However the term $(1 + \epsilon \sin t) \sin \theta W_0^2/2$ will generate a contribution to $V_{14}$, which will have a non-zero mean since the product $W_0^2$ contains time-independent terms. It is easily seen that, since $W_0 = O(W_2^2)$, $V_{14} = O(W_2^2)$. This is the scale of the secondary velocity that is generated in the Stokes layer by the centrifugal force effects that arise because the tube is curved. ($V_{14}$ is just Lyne’s quasi-steady secondary velocity.)

The solutions of (2.27) and (2.29) do not match to the core solutions as $\eta \to \infty$. The solutions for both $V_{12}$ and $V_{14}$ satisfy the no-slip condition at the boundary, but, in both cases, the velocity does not fall to zero as $\eta \to \infty$. Instead, these secondary motions ‘drag’ the fluid in the core of the tube, thus generating secondary motions outside the boundary layer. For non-zero $\Phi$, both $V_{12}$ and $V_{14}$ will generate steady secondary streaming in the core.

Putting these results back into dimensional form, we obtain

$$\hat{V}_{12x} = O(\nu \epsilon W_p(2\delta_0)/a) \quad (2.30)$$

and

$$\hat{V}_{14x} = O(\nu W_p^2(2\delta_0)/(ax^2)) \quad (2.31)$$

where $\hat{V}_{12x}$ and $\hat{V}_{14x}$ are the velocities at the edge of the Stokes layer. Thus the secondary streaming Reynolds numbers are defined to be

$$R_{s_1} = \hat{V}_{12x} a/\nu = \epsilon W_p(2\delta_0),$$

$$2R_s = \hat{V}_{14x} a/\nu = W_p^2(2\delta_0)/ax^2. \quad (2.32)$$

$R_s$ is the secondary streaming Reynolds number based on the scale of the secondary velocity that arises because the tube is curved, and is the same as that determined by Lyne for a fixed curved tube. $R_{s_1}$ is the secondary streaming Reynolds number based on the scale of the secondary velocity that arises from the time-dependence of curvature.

Substitution of $\delta, \dot{\delta}$ and $\ddot{\delta}$ into equations (2.3a–d) reveals that the product of $\alpha \epsilon$ frequently occurs. Thus we define a new parameter, $R_s$, as

$$R_s = \alpha \epsilon = R_{s_1}/[(2R_s)^{1/2}(2\delta_0)^{1/2}]. \quad (2.33)$$
2.3. Parameter scalings

Equations (2.3a–d) are very complicated and cannot be solved as they stand. In order to simplify the problem and to isolate the effects of the various new terms upon the magnitude and distribution of wall shear stress, and the generation of steady secondary streaming, various perturbative techniques are carried out.

The four governing parameters of the problem are \( \delta_0, R_t, (2R_s)^{1/2} \) and \( z \). In this paper we examine one particular choice of parameter scalings. Initially \( z \) and \( R_s \) are chosen to be order one compared with \( \delta_0 \) and \( R_t \), although they are later taken to be either large or small. We shall consider \( \delta_0 \) to be the smallest parameter of the problem and take \( \delta_0 \to 0 \), while keeping \( R_t = O(1) \). \( R_t \) is then taken to be small because \( \epsilon \ll 1 \). (\( \epsilon \) is chosen to be sufficiently small that \( R_t \) can be taken to be small even in the case where \( z \gg 1 \), as in §5.) We shall show that the leading-order flow is just that driven by an oscillatory pressure gradient in a fixed curved pipe (§3). At \( O(R_t) \), effects due to the unsteadiness of the curvature are first seen. The details of this problem are described in §§4, 5 for low and high frequencies respectively.

3. Perturbation to oscillatory flow in a fixed, curved pipe

We now proceed to solve equations (2.3a–d) together with the boundary condition (2.5) for flows driven by both low- and high-frequency pressure gradients subject to the parameter scaling discussed in §2.3.

The axial velocity is rescaled so that the centrifugal force terms in the governing equations are formally the same order of magnitude as the viscous and inertial terms. Thus we take \( w \to w/(2\delta_0)^{1/2} \): we also take \( s \to s/(2\delta_0)^{1/2} \). \( \delta, \dot{\delta} \) and \( \ddot{\delta} \) are substituted into (2.3a–d) and \( R_t \) is defined as in (2.33). Note that \( -\partial p(s)/\partial s = \alpha^3(2R_s)^{1/2} \cos(t + \Phi)/(2\delta_0)^{1/2} \). Since \( \delta_0 \) is the smallest parameter of the problem, we delete all terms that are \( O(\delta_0) \) relative to the leading terms. Next a small-parameter expansion is carried out in powers of \( R_t \):

\[
\begin{align*}
\bar{w} & = \bar{w}_0 + R_t \bar{w}_1 + \cdots, \\
\bar{u} & = \bar{u}_0 + R_t \bar{u}_1 + \cdots, \\
\bar{v} & = \bar{v}_0 + R_t \bar{v}_1 + \cdots, \\
\bar{p} & = \bar{p}_0 + R_t \bar{p}_1 + \cdots,
\end{align*}
\]

(3.1)

and we consider the equations at various orders of \( R_t \). At leading order, \( O(R_t^0) \), the equations are the same as for flow in a fixed curved pipe.

At \( O(R_t) \) the equations are (dropping overbars):

continuity:

\[ u_{1r} + \frac{1}{r} v_{1\theta} + w_{1s} + \frac{1}{r} u_1 = 0, \]

(3.2a)

e_r:

\[
x^2 u_{1r} + u_0 u_{1r} + u_1 u_{0r} + \frac{1}{r} v_0 u_{1\theta} + \frac{1}{r} v_1 u_{0\theta} + w_0 u_{1s} + w_1 u_{0s} - \frac{2}{r} v_0 v_1 \\
+ \frac{1}{4} x^3 \cos \theta s^2 \sin t - x \cos \theta \cos t w_0 - \cos \theta w_0 w_1 - \frac{1}{x} \cos \theta \sin t w_0^2 \\
= -\bar{p}_{1r} - \frac{1}{r} \left\{ \frac{\partial}{\partial \theta} \left[ \frac{1}{r} (v_1 + rv_{1r} - u_{1\theta}) \right] - \frac{\partial}{\partial s} \left(-rv_{1s}\right) \right\},
\]

(3.2b)
Oscillatory flow in a tube of time-dependent curvature. Part 1

$e_\theta$:

\[ x^2 v_{1r} + u_0 v_{1r} + u_1 v_{0r} + \frac{1}{r} v_0 v_{1r} + \frac{1}{r} v_1 v_{0r} + w_0 v_1 + w_1 v_0 + \frac{1}{r} (u_0 v_1 + u_1 v_0) \]

\[ - \frac{1}{2} x^3 \sin \theta s^2 \sin t + \alpha \sin \theta \cos t w_0 + \sin \theta w_0 w_1 + \frac{1}{x} \sin \theta \sin t w_0^2 \]

\[ = -\frac{1}{r} p_{1\theta} + \left\{ \frac{\partial}{\partial r} \left[ \frac{1}{r} (v_1 + r v_{1r} - u_{1r}) \right] - \frac{\partial}{\partial s} \left( \frac{1}{r} w_{1\theta} \right) \right\}, \quad (3.2c) \]

$e_\zeta$:

\[ x^2 w_1 + u_0 w_1 + u_1 w_{0r} + \frac{1}{r} (v_0 w_{1r} + v_1 w_{0r}) + w_0 w_1 + w_1 w_0 = \nabla^2 w_1, \quad (3.2d) \]

Thus at this order, effects due to the time-dependence of curvature are present. Note that new terms exist in equations (3.2b,c) but there are no additional new terms to the axial momentum equation (3.2d). By assuming $R_t \ll 1$ we have taken $R_s \ll (2 R_s)^{1/2}$ (see equation (2.33)) so that the velocity at the edge of the Stokes layer that arises due to the time-dependence of curvature is less than that arising because the tube is curved.

4. Low frequency

4.1. Solution

In this section we consider flows at low frequency, $\alpha \ll 1$ (in fact $\delta_0 \ll R_t \ll \alpha \ll 1$). We therefore consider a small-parameter expansion in powers of $\alpha$. $R_s$ is assumed to remain $O(1)$ compared with the other governing parameters.

At $O(R^0)$, the equations for flow in a fixed curved pipe are obtained (Mullin & Greated 1980b). The radial and azimuthal equations may be combined in a single equation for the streamfunction, $\psi_0$. The following expansions for the axial velocity, $w_0$, and the streamfunction, $\psi_0$, are then considered:

\[ w_0 = x^3 w_{00} + x^4 w_{01} + \cdots, \]
\[ \psi_0 = x^6 \psi_{00} + x^7 \psi_{01} + \cdots. \quad (4.1) \]

The leading-order axial velocity, driven by the applied oscillatory pressure gradient, is just

\[ w_{00} = \frac{1}{3} (2 R_s)^{1/2} (1 - r^2) \cos (t + \Phi). \quad (4.2) \]

Thus the leading-order axial wall shear stress is

\[ -w_{000r} |_{r=1} = \frac{1}{3} (2 R_s)^{1/2} \cos (t + \Phi). \quad (4.3) \]

The balance of the local inertia term with the viscous term in the axial velocity equation leads to the following further corrections to the axial velocity:

\[ \begin{aligned} w_{0t} &= 0, \\
 w_{03} &= \frac{1}{8} (2 R_s)^{1/2} (3 - r^2) (1 - r^2) \sin (t + \Phi), \\
 w_{03} &= 0, \\
 w_{04} &= \frac{1}{8} (2 R_s)^{1/2} \left( \frac{15}{4} + \frac{5}{4} r^2 - \frac{1}{4} r^4 + \frac{1}{4} r^6 \right) \cos (t + \Phi). \end{aligned} \quad (4.4) \]

Both $w_{02}$ and $w_{04}$ are symmetric in $r$. Note that we let $w_{04} = w_{04}^k + w_{04}^s$ since the axial velocity at this order consists of two parts: a symmetric part, $w_{04}^k$, and a term, $w_{04}^s$. 
that is asymmetric in \( r \) (see below). The leading-order streamfunction is

\[
\psi_{00} = \frac{1}{8 \times 144} R_s r (1 - r^2)^2 (1 - \frac{1}{4} r^2) \sin \theta \cos^2(t + \Phi).
\]  

(4.5)

The leading-order azimuthal wall shear stress, (2.9b), is

\[
\psi_{00r} |_{r=1} = \frac{1}{172} R_s \sin \theta \cos^2(t + \Phi).
\]  

(4.6)

Equations (4.2) and (4.5) are the quasi-steady equivalent of the leading-order Poiseuille axial velocity and streamfunction found for Dean flow. The term \( (1/r) \psi_{00\theta} \), proportional to the radial velocity, is positive on \( \theta = 0 \) (and negative on \( \theta = \pi \)). Thus, for small frequencies, the steady component of streaming in the core proceeds across the tube from the inside of the bend to the outside (as in Dean flow). This secondary streamfunction advects the leading-order quasi-steady Poiseuille velocity, \( w_{00} \), which results in a ‘skewing’ of the previously symmetric axial velocity profile. This axial velocity perturbation is

\[
w_{04} = \frac{1}{45 \times 85} (2 R_s)^{3/2} r (1 - r^2) (19 - 21 r^2 + 9 r^4 - r^6) \cos \theta \cos^3(t + \Phi).
\]  

(4.7)

Thus the curvature of the pipe leads to the distortion of the axial velocity profile and an increase in the wall shear stress on the outside of the wall. The results are \( s \)-independent. At this order there are no new effects due to the unsteadiness of the curvature. In order to understand the effects of time-dependent curvature on the flow, we must consider the equations at \( O(R_t) \).

To solve the \( O(R_t) \) equations, (3.2a–d), we consider the following small-parameter expansion in powers of \( \alpha \):

\[
\begin{align*}
    w_1 &= \alpha^7 w_{10} + \alpha^8 w_{11} + \cdots, \\
    u_1 &= \alpha^4 u_{10} + \alpha^5 u_{11} + \cdots, \\
    v_1 &= \alpha^4 v_{10} + \alpha^5 v_{11} + \cdots, \\
    p_1 &= \alpha^3 p_{10} + \alpha^4 p_{11} + \cdots.
\end{align*}
\]  

(4.8)

(The expansion in powers of \( \alpha \) is taken to a sufficiently high degree to obtain the first new effects from the time-dependent curvature.)

At \( O(\alpha^4) \) equations (3.2b,c) give

\[
p_{10} = -\frac{1}{4} r \cos \theta s^2 \sin t;
\]  

(4.9)

the fluid pressure adjusts to balance the new centrifugal-force term that arises from the time-dependence of pipe curvature. It prevents the fluid being ‘flung’ out along the pipe.

At \( O(\alpha^3) \) and \( O(\alpha^5) \) equation (3.2a) admits the existence of streamfunctions \( \psi_{10} \) and \( \psi_{11} \) such that \( u_{10} = (1/r) \psi_{10\theta} \), \( v_{10} = -\psi_{10r} \) and similarly for \( u_{11} \) and \( v_{11} \).

At \( O(\alpha^5) \) equations (3.2b,c) can be combined to give

\[
\nabla^4 \psi_{10} = \frac{1}{2} (2 R_s)^{1/2} r \sin \theta s \cos t \cos (t + \Phi),
\]  

(4.10)

which has solution

\[
\psi_{10} = \frac{1}{2} \times 384 (2 R_s)^{1/2} r (1 - r^2)^2 \sin \theta s (\cos (2t + \Phi) + \cos \Phi).
\]  

(4.11)

This secondary streaming results from the interaction between the viscous terms and the Coriolis force term that arises from the time-dependence of curvature.
At leading order (3.2$d$) gives an equation for $w_{10}$ which has solution

$$w_{10} = \frac{1}{1832} (2R_s) \cos \theta s \cos t \cos^2(t + \Phi) r(1 - r^2)(3 - 3r^2 + r^4). \quad (4.12)$$

4.2. Discussion

The first new contribution to the axial velocity from the time-dependence of curvature is $w_{10}$. It is an $O(R_t)$ perturbation to the axial velocity profile obtained for a fixed curved pipe, and is due to the advection of the leading-order quasi-steady Poiseuille velocity by the new secondary streaming which results in a ‘skewing’ of the axial velocity profile. This skewing may enhance or diminish the skewing of the axial velocity profile that arises merely due to the pipe being curved.

At any given time $t$, the direction of the secondary streaming and hence the ‘skewing’ of the axial velocity depends on the relative signs of $\cos t$ and $\cos (t + \Phi)$ (since $r(1 - r^2)^2$ and $r(1 - r^2)(3 - 3r^2 + r^4)$ are always positive for $r$ between 0 and 1). For example, if $\cos (t + \Phi)$ is positive so that the flow is in a positive sense along the pipe, and $\cos t > 0$, so that the pipe is rotating in a negative sense, the streaming is from the inside of the bend to the outside. Thus the skewing of the axial velocity will be towards the outside wall of the pipe, enhancing the effect of pipe curvature and increasing the wall shear stress there.

If, however, $\cos t < 0$ while $\cos (t + \Phi) > 0$, the streaming is in the opposite sense and the skewing of the axial velocity profile will be towards the inside wall, increasing the wall shear stress there. The dependence of $\psi_{10}$ on the axial coordinate $s$ means that its influence on the axial velocity profile will become more pronounced as we move away from the fixed origin (figure 2). The new streaming, $\psi_{10}$, is an order of magnitude smaller than the streaming that results solely from the tube curvature and thus its influence on the axial velocity profile will be small.

We are particularly interested in the properties of the steady components of the secondary streaming and the wall shear stress. The direction of the steady component of the $s$-dependent secondary streaming, $\psi_{10}$, depends on the sign of $\cos \Phi$. If $\cos \Phi$ is positive then the steady part of $(1/r)\psi_{10}$ is positive on $\theta = 0$ (and negative on $\theta = \pi$) and thus the steady component of the secondary streaming proceeds across the centre...
of the pipe from the inside of the bend to the outside, in the same sense as for steady curvature. However, if \( \cos \Phi \) is negative then the steady secondary streaming is in the opposite sense. When \( \cos \Phi = 0 \), i.e. when the time-dependent curvature variations are out of phase with the oscillatory pressure gradient by \( \pi/2 \), there is no contribution to the steady secondary streaming at this order.

The axial wall shear stress perturbation, due to the advection of \( w_{00} \) by the secondary streaming, \( \psi_{10} \), is

\[
-w_{10r}|_{r=1} = \frac{1}{2\Omega} (2R_s) \cos \theta \cos t \cos^2(t + \Phi),
\]

and thus, when \( \cos t > 0 \), the wall shear stress is increased on the outside wall and decreased on the inside wall (with the opposite being the case when \( \cos t < 0 \)). The azimuthal wall shear stress perturbation, from (4.11), is

\[
\psi_{10r}|_{r=1} = \frac{1}{\Omega} (2R_s)^{1/2} \sin \theta \cos (2t + \Phi) + \cos \Phi.
\]

In the above analysis \( R_s \) is assumed to remain \( O(1) \) as all the other parameters tend to zero.

The time-dependency of the secondary streaming, \( \psi_{10} \), means that \( w_{10} \) may be very small for certain ranges of \( t \). Within this range, the axial velocity terms at higher powers of \( s \) may have an appreciable effect.

5. High frequency

5.1. Secondary core flow equations

In this section, the small-\( Rt \) perturbation to Lyne’s solution for flow driven by a high-frequency oscillatory pressure gradient in a fixed curved pipe is found. We assume that \( \delta_0 \ll R_t \ll 1/\alpha^2 \ll 1 \) and \( R_s \) remains \( O(1) \) initially.

At \( O(R_t^0) \) the equations are the \( s \)-independent equations for flow in a fixed, curved pipe. Lyne (1971) analysed this problem and showed that the leading-order steady streaming in the core \( \psi_{00}^{(s)} = R_s \chi_{00} \) is given by the solution to

\[
\begin{align*}
\frac{1}{R_s} \nabla^4 \chi_{00} + \frac{1}{r} \left( \chi_{00} \frac{\partial}{\partial \theta} - \chi_{00} \frac{\partial}{\partial r} \right) \nabla^2 \chi_{00} &= 0, \\
\chi_{00} &= 0, \quad \chi_{00r} = \frac{1}{4} \sin \theta \quad \text{on} \quad r = 1,
\end{align*}
\]

which makes it clear that \( R_t \) is the steady streaming Reynolds number (see §2.2).

As found in §4 for small \( \alpha \), at leading order the equations contain no new terms due to the unsteadiness of curvature. Thus it is necessary to consider the equations at \( O(R_s) \), (3.2a–d), which we have to consider the equations at \( O(R_s) \), (3.2a–d), which we have to consider the equations at \( O(R_s) \), (3.2a–d), which suggest that \( u_1, v_1, w_1 \) and \( p_1 \) depend on \( s \) as follows:

\[
\begin{align*}
u_1 &= su_{10} + u_{11}, \\
v_1 &= sv_{10} + v_{11}, \\
w_1 &= w_{10}, \\
p_1 &= s^2 p_{10} + sp_{11} + p_{12}.
\end{align*}
\]

We are most interested in finding the \( s \)-dependent velocities since it is these that arise from the introduction of time-dependent curvature. We find at \( O(s^2) \) that

\[
p_{10} = -\frac{1}{4} z' r \cos \theta \sin t,
\]

which is just the fluid pressure adjustment to balance the new centrifugal-force terms.
The s-dependent equations are:

continuity

\[ u_{10r} + \frac{1}{r} v_{100} + \frac{1}{r} u_{10} \theta = 0, \]

\[ e_r \]

\[ \alpha^2 u_{10r} + u_0 u_{10r} + u_{10} u_{0r} + \frac{1}{r} v_0 u_{100} + \frac{1}{r} v_{10} u_{00} - \frac{2}{r} v_0 v_{10} - \alpha \cos \theta \cos t w_0 \]

\[ = -p_{11r} - \frac{1}{r} \frac{\partial}{\partial \theta} \left[ \frac{1}{r} (v_{10} + rv_{10} - u_{10}) \right], \]

\[ e_\theta \]

\[ \alpha^2 v_{10r} + u_0 v_{10r} + u_{10} v_{0r} + \frac{1}{r} v_0 v_{100} + \frac{1}{r} v_{10} v_{00} + \frac{1}{r} (u_0 v_{10} + u_{10} v_0) + \alpha \sin \theta \cos t w_0 \]

\[ = -\frac{1}{r} p_{110} + \frac{\partial}{\partial r} \left[ \frac{1}{r} (v_{10} + rv_{10} - u_{10}) \right], \]

From (5.4a) we know that there exists a streamfunction \( \psi_{10} \) such that

\[ u_{10} = \frac{1}{r} \psi_{10} \theta, \]

\[ v_{10} = -\frac{1}{r} \psi_{10} r. \]

Thus (5.4b) and (5.4c) may be combined to give

\[ \alpha^2 \nabla^2 \psi_{10} - \alpha \cos t \left( \frac{1}{r} \cos \theta w_{00} + \sin \theta w_{0r} \right) \]

\[ = \nabla^4 \psi_{10} + \frac{1}{r} \left( \psi_{00} \frac{\partial}{\partial \theta} - \psi_{00} \frac{\partial}{\partial r} \right) \nabla^2 \psi_{10} + \frac{1}{r} \left( \psi_{10} \frac{\partial}{\partial \theta} - \psi_{10} \frac{\partial}{\partial r} \right) \nabla^2 \psi_{0}. \]

We are interested in determining the secondary motions generated in the core; however, it is first necessary to determine the behaviour of the flow in the boundary layer, since it is that which drives the secondary motions in the core.

In the viscous boundary layer we take

\[ \Psi_{10} = \frac{1}{\alpha} \left( \Psi_{100} + \frac{1}{\alpha} \Psi_{101} + \ldots \right) \]

and the boundary layer coordinate is again

\[ \eta = \frac{\alpha}{\sqrt{2}} (1 - r). \]

The equation for \( \Psi_{100} \) is, from (5.5)

\[ \frac{1}{2} \Psi_{100 \eta \eta \eta} - \Psi_{100 \eta \eta} = \sqrt{2} \cos t \sin \theta W_{00 \eta} \]

where \( W_{00} = (2R_s)^{1/2} (\sin (t + \Phi) + e^{-\eta} \sin (t + \Phi - \eta)) \) (see (2.14)), which has solution

\[ \Psi_{100} = \sin \theta (2R_s)^{1/2} \left[ \frac{1}{2} e^{-\eta} \sin (2t + \Phi - \eta + \frac{1}{4} \pi) + \frac{1}{2} e^{-\eta} \sin (-\Phi + \eta - \frac{1}{4} \pi) \right. \]

\[ - \frac{1}{2} \frac{1}{\sqrt{2}} e^{-\sqrt{2} \eta} \sin (2t + \Phi - \sqrt{2} \eta + \frac{1}{4} \pi) - \frac{1}{\sqrt{2}} \eta \cos \Phi \]

\[ + \frac{1}{2} \sin (\Phi + \frac{1}{4} \pi) + \frac{1}{2} \sqrt{2} \left( 1 - \sqrt{2} \right) \sin (2t + \Phi + \frac{1}{4} \pi) \].

\( \Psi_{100} \) contains time-independent terms, because although \( W_{00 \eta} \) has no mean part, the product \( \cos t W_{00 \eta} \) does have a mean part. Hence we see that, as \( \eta \to \infty \), not only
does $\Psi_{100}$ not tend to zero (displacement effect) but also

$$\Psi_{100} \to -\frac{1}{2} (2R_s)^{1/2} \cos \Phi \sin \theta.$$  \hspace{1cm} (5.10)

After an analysis similar to Lyne’s the equation for the steady streaming in the core $\psi_{100}$, derived from (5.5), is given by the solution to

$$\nabla^4 \psi_{100} + \frac{1}{r} \left( \psi_{100r} \frac{\partial}{\partial r} - \psi_{100\theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \psi_{100} = 0,$$  \hspace{1cm} (5.11)

subject to the boundary conditions derived from (5.9) and (5.10) that

$$\psi_{100} = 0, \quad \psi_{100r} = \frac{1}{2} (2R_s)^{1/2} \sin \theta \cos \Phi \quad \text{on} \quad r = 1.$$  \hspace{1cm} (5.12)

(Here $\psi_{00}$ and $\psi_{100}$ are the leading-order steady terms of a power series expansion in $\alpha$ for $\psi_0$ and $\psi_{10}$.)

If we make the transformations $\psi_{100} = R_s^{1/2} \chi_{100}$ and $\psi_{00} = R_s \chi_{00}$, then the equation to solve is

$$\frac{1}{R_s} \nabla^4 \chi_{100} + \frac{1}{r} \left( \chi_{100r} \frac{\partial}{\partial r} - \chi_{100\theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \chi_{100} = 0,$$  \hspace{1cm} (5.13)

subject to the boundary conditions:

$$\chi_{100} = 0, \quad \chi_{100r} = \frac{1}{2} \sqrt{2} \sin \theta \cos \Phi \quad \text{on} \quad r = 1.$$  \hspace{1cm} (5.14)

We will clearly be able to solve (5.13) for $\chi_{100}$ using the same methods as Lyne (1971) used for (5.1).

5.2. Analytical streaming solution at small $R_s$

Analytic progress towards a solution of (5.1) and (5.13) may be made in the limit of small $R_s$ by considering a small-parameter expansion in powers of $R_s$ (Lyne 1971):

$$\chi_{00} = \chi_{000} + R_s \chi_{001} + \cdots, \quad \chi_{100} = \chi_{1000} + R_s \chi_{1001} + \cdots.$$  \hspace{1cm} (5.15)

The leading-order steady streaming problem (5.1) is reduced to a simple Stokes flow problem, with solution

$$\chi_{00} = -\frac{1}{8} (1 - r^2) \sin \theta - R_s \frac{1}{1072} r^2 (1 - r^2)^2 \sin 2\theta + O(R_s^2).$$  \hspace{1cm} (5.16)

The steady secondary streaming in the core, to leading order in $R_s$, proceeds across the tube from the outside of the bend to the inside. This is in the opposite direction to the secondary flow in a pipe when the axial pressure gradient is steady (or when $\alpha \ll 1$ as in the quasi-steady case detailed in §4). Thus whereas the intuitive idea of ‘outwards centrifuging’ is valid for steady flow, it is not relevant for oscillatory flow and the apparent ‘centrifuging’ is negative and thus directed inwards. Lyne verified this analytic result experimentally.

The leading-order solution of (5.13), subject to the boundary conditions (5.14), is

$$\chi_{1000} = -\frac{1}{4} \sqrt{2} (1 - r^2) \sin \theta \cos \Phi.$$  \hspace{1cm} (5.17)

Thus when $\cos \Phi > 0$ this contribution to the streaming proceeds across the tube from the outside of the bend to the inside (i.e. in the same direction as Lyne’s) and thus enhances the streaming arising due to the tube being curved. However, for $\cos \Phi < 0$ the contribution is in the opposite direction, and thus diminishes the streaming. The dependence of the steady streaming on $s$, given by (5.2), means that
its magnitude will increase with distance from the fixed origin. Of course, since we have considered a small-\(R_t\) parameter expansion the effect of this perturbation on the leading-order secondary streaming will be small.

5.3. Numerical solution at intermediate values of \(R_s\)

In order to determine the behaviour of the secondary streaming in the core at intermediate values of the secondary streaming Reynolds number \(R_s\) we must solve equations (5.1) and (5.13) numerically, which we do using a finite difference method. Haddon (1982) has solved Lyne’s problem numerically, also using a finite difference method, but with a different iteration procedure.

5.3.1. Numerical method

Equations (5.1) and (5.13) must be linearized in order to obtain a system of linear algebraic equations, which can then be solved numerically. This is done using the Newton–Raphson scheme.

Let \(n\) label the iteration number, and define \(\xi_{00}^n = \chi_{00}^n - \chi_{00}^{n-1}\).

The Newton–Raphson iterative scheme for (5.1) is

\[
\nabla^4 \xi_{00}^n + R_s \frac{1}{r} \left( \frac{\partial \chi_{00}^{n-1}}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \chi_{00}^{n-1}}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \xi_{00}^n + R_s \frac{1}{r} \left( \frac{\partial \xi_{00}^n}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \xi_{00}^n}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \chi_{00}^{n-1} = -\nabla^4 \chi_{00}^{n-1} - R_s \frac{1}{r} \left( \frac{\partial \chi_{00}^{n-1}}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \chi_{00}^{n-1}}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \chi_{00}^{n-1},
\]

which is iterated until a converged solution for \(\xi_{00}^n\) is obtained. Similarly, by defining \(\xi_{100}^n = \chi_{100}^n - \chi_{100}^{n-1}\),

\[
\nabla^4 \xi_{100}^n - R_s \frac{1}{r} \left( \frac{\partial \chi_{100}^{n-1}}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \chi_{100}^{n-1}}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \xi_{100}^n - R_s \frac{1}{r} \left( \frac{\partial \xi_{100}^n}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \xi_{100}^n}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \chi_{100}^{n-1} = -\nabla^4 \chi_{100}^{n-1} - R_s \frac{1}{r} \left( \frac{\partial \chi_{100}^{n-1}}{\partial r} \frac{\partial}{\partial r} - \frac{\partial \chi_{100}^{n-1}}{\partial \theta} \frac{\partial}{\partial \theta} \right) \nabla^2 \chi_{100}^{n-1},
\]

We wish to solve the flow in the circular plane of cross-section of the pipe. However, since the flow is symmetric about the diameter \(\theta = 0, \pi\) we need only compute the flow in one half of the pipe. The computational domain is therefore a semi-circle, polar coordinates \((r, \theta)\) are used and grid lines are chosen along lines of constant radius and constant \(\theta\).

Both \(\chi\) and \(\partial \chi/\partial r\) are given on \(r = 1\) ((5.1) and (5.14)). Since the boundary of the semi-circular domain is a streamline we have that \(\chi = 0\) on \(\theta = 0, \pi\). We also have zero normal velocity (in the \(\theta\)-direction) on the line \(\theta = 0, \pi\). Thus \(\partial u/\partial \theta = (1/r)(\partial^2 \chi/\partial \theta^2) = 0\) on \(\theta = 0, \pi\).

The governing equations (5.19) and (5.21) are discretized using a second-order-accurate finite-difference technique. A set of linear algebraic equations is obtained, which can be written in matrix form. At each Newton–Raphson iterative step the linear matrix equations are solved using the alternating-direction-implicit (ADI) relaxation method.
method. In order to make the ADI scheme converge, it is necessary to under-relax, using a relaxation parameter, $c$, where $0 < c < 1$. The value of $c$ necessary for convergence depends on the size of the mesh used. In general, the finer the mesh, the smaller the value of $c$ required to obtain convergence.

The streamfunction for a given value of $R_\epsilon$ is found using the Newton–Raphson method as follows. Suppose $\Delta R_\epsilon$ is an increment in $R_\epsilon$ and $\chi_L$ is the solution to the linear problem, $\nabla^2 \chi = 0$, corresponding to $R_\epsilon = 0$. Then the basic stages of the numerical algorithm are:

(i) Solve the streamfunction equation (5.19), with the corresponding boundary conditions, at $R_\epsilon = \Delta R_\epsilon$ using $\chi = \chi_L$ as an initial guess to the solution.

(ii) Update $R_\epsilon$ to $R_\epsilon = 2\Delta R_\epsilon$, and solve for the streamfunction using the converged solution for the streamfunction at $R_\epsilon = \Delta R_\epsilon$, as an initial guess.

(iii) Repeat until the streamfunction at the required value of $R_\epsilon$ is obtained.

In order to check the accuracy of the numerical code, we examined the effect of tightening the tolerance criteria for the ADI and Newton–Raphson schemes ($\epsilon_1$ and $\epsilon_2$ respectively) and refining the mesh. To ensure that the chosen tolerances were sufficiently small, we computed the solution at tighter tolerance values and investigated how the obtained numerical solution changed both quantitatively and qualitatively. This was done for a range of values of $R_\epsilon$, but most importantly for the highest value of $R_\epsilon$ at which results were obtained ($R_\epsilon = 300$) since this case will be the most sensitive. For $R_\epsilon = 300$, initially we took $\epsilon_1 = \epsilon_2 = 10^{-5}$. We then computed the solutions at $\epsilon_1 = \epsilon_2 = 10^{-6}$. It was found that the maximum difference between the values of $\psi_{i,j}$ obtained at the two different tolerances was $O(10^{-5})$, and the qualitative behaviour of the obtained solutions was unchanged.

The effect of grid-refinement on the numerical solution was also investigated. Initially, the solutions at various $R_\epsilon$ were calculated on the coarsest mesh possible that allowed a convergent solution to be found (as the value of $R_\epsilon$ increases, it is necessary to refine the mesh in order to converge to a solution.) For example, for values of $R_\epsilon$ up to about 170, solutions were obtained using a $(21 \times 21)$ grid, whereas for values of $R_\epsilon$ between 170 and 300, solutions were obtained on a $(31 \times 31)$ mesh. For $R_\epsilon = 10$ and $R_\epsilon = 50$, we also calculated the solution on grids of $(41 \times 41)$ and $(81 \times 81)$. For $R_\epsilon = 10$, the maximum difference between the solutions found on a $(21 \times 21)$ and a $(41 \times 41)$ grid was order $10^{-4}$ and on a $(41 \times 41)$ and $(81 \times 81)$ grid was order $10^{-5}$. For $R_\epsilon = 50$ the difference was $O(10^{-4})$. Grid refinements were also carried out for $R_\epsilon = 100$ where again the differences were $O(10^{-4})$. For $R_\epsilon = 200$ and 300 it was found that the differences were $O(10^{-5})$. Hence the underlying physics of the system was captured using the computationally less expensive coarser mesh.

We looked at the effect of tightening the tolerances and refining the mesh on the maximum value of vorticity, $\omega$, that is obtained. Since vorticity involves second-order derivatives, $\omega = -\nabla^2 \psi$, it is a good quantity to consider, since it is likely to be the most poorly behaved of the system. It was found that making the tolerances tighter had a small effect on the maximum value of vorticity obtained. For example, for $R_\epsilon = 10$, the difference between solutions obtained at the tolerances $\epsilon_1 = \epsilon_2 = 10^{-6}$ and $\epsilon_1 = \epsilon_2 = 10^{-7}$ was 0.04%. For $R_\epsilon = 300$, the difference between solutions obtained at $\epsilon_1 = \epsilon_2 = 10^{-5}$ and $\epsilon_1 = \epsilon_2 = 10^{-6}$ was 0.3%.

Also, for $R_\epsilon = 10$, halving the mesh size resulted in a 0.02% change in the value of the maximum vorticity. For $R_\epsilon = 50$, halving the mesh size resulted in a 0.1% change in the maximum value of the vorticity. For $R_\epsilon = 100$, the change was again 0.1%. As the value of $R_\epsilon$ increased, however, the percentage change also increased. For $R_\epsilon = 200$, it was found that refining the mesh resulted in a 6% change and for
Oscillatory flow in a tube of time-dependent curvature. Part 1

Figure 3. Secondary streamfunction, $\chi_{00}$, for $R_s = 0$.

Figure 4. Numerical solution for the streamfunction, $\chi_{00}$, for increasing $R_s$:
(a) $R_s = 50$; (b) $R_s = 250$.

$R_s = 300$, the change was 8%. The computational scheme would clearly have to be refined if we were to go to higher values of $R_s$.

5.3.2. Results

The leading-order secondary streamfunction, $\chi_{00}$, at $R_s = 0$ is shown in figure 3. The streamfunction is symmetric about the line $\theta = \pi/2$. Figure 4 shows how the flow develops with increasing secondary streaming Reynolds number, $R_s$. The secondary streaming still proceeds across the centre of the tube from the outside of the bend towards the inside. As the value of $R_s$ increases, the advection terms in (5.1) become increasingly important, and it is these effects that we observe here. Figure 4(a) shows the secondary streamlines when $R_s = 50$. The stationary point, or vortex centre, is shifted in the direction of the driving secondary flow, towards the outer wall of the pipe. As the value of $R_s$ increases to $R_s = 250$, the vortex centre continues to be shifted in the direction of the flow (figure 4b).

We note that the solutions for the secondary streamfunction, $\chi_{00}$, at $R_s = 50$ and $R_s = 250$ look similar (figure 4a,b). As $R_s$ increases, however, there is a definite change in the streamfunction solution which may be clearly seen by considering the difference between the solutions at various $R_s$. Figure 5(a) shows the difference between the streamfunction values corresponding to $R_s = 50$ and $R_s = 10$. The convective motion consists approximately of two counter-rotating eddies. This is consistent with the analytic solution found at small values of $R_s$, which gives the first correction to the viscous Stokes layer solution to be proportional to $\sin 2\theta$. As the value of $R_s$ increases, the convective motion still has the form of two counter-rotating eddies, but now they are no longer symmetrical and, moreover, are of unequal size (figure 5b).
Next we consider the perturbation to the secondary streaming that arises due to the introduction of time-dependent curvature. We present results for the secondary streamfunction, $\chi_{100}$, at values of $R_s = 10$ and 100 (figure 6a,b).

The interesting feature here is that the direction of the perturbation to the steady secondary flow is dependent on the phase difference, $\Phi$, between the oscillatory pressure gradient and the time-dependent curvature. When $\cos \Phi$ is positive, the perturbation to the secondary flow is in the same direction as for Lye. When $\cos \Phi$ is negative the reverse is true. The results presented in figure 6(a,b) are for $\cos \Phi = 1$.

At this order the advection effects are stronger. Whereas in (5.1) we saw that the streamfunction $\chi_{00}$ was advected by itself, here the streamfunction $\chi_{100}$ is advected by $\chi_{00}$ and $\chi_{00}$ is advected by $\chi_{100}$ (see equation (5.13)). Thus the advective effects are doubled, and this is reflected in the greater value of the streamfunction $\chi_{100}$ for a given value of $R_s$, and the greater advection of the vorticity for a given value of $R_s$.

The direction in which the vortex centre of the secondary streamlines is 'skewed' is independent of the sign of $\cos \Phi$. It is always towards the outside wall for the range of secondary streaming Reynolds numbers that we consider here. Thus, when $\cos \Phi$ is negative, we have the interesting situation that the direction of skewing of the stagnation point of the secondary streamlines is in the opposite direction to the driving circumferential velocity. When $\cos \Phi = 0$, $\chi_{100}$ is identically zero and there is no contribution to the steady secondary streamfunction at this order.

The dependence of $\chi_{100}$ on $s$ means that the magnitude of the perturbation to the streamfunction will increase with distance from the fixed origin.

5.4. Solution at asymptotically large values of $R_s$

At large values of $R_s$ it is hypothesized that the effect of viscosity is confined to boundary layers of thickness $R_s^{-1/2}$ near $r = 1$. These come together at $\theta = 0$ and
erupt across the centre of the tube. Lyne (1971) postulated a scheme in which the wall boundary layers are supplemented by a layer of thickness $R - 1/2$ across the plane of symmetry, the bulk of the core flow being inviscid. In the inviscid region the streamlines are closed, and therefore the vorticity is uniform and of opposite sign in the two regions (Batchelor 1956); one reason why the shear layer across the middle of the tube is required is to smooth out the discontinuity in vorticity between the two inviscid regions.

Haddon (1982) gave a numerical solution of Lyne’s (1971) steady streaming problem (equation (5.1) above) which he took to large values of $R_s$ (up to 3000). He was able to show that the vorticity distribution in one half of the pipe cross-section (along $r = 1/2$, $0 < \theta < \pi$, and along $\theta = \pi/2$, $0 < r < 1$) became approximately uniform over a substantial part of the region. Moreover, the value of the vorticity, in the limit as $R_s \to \infty$ and the mesh size $\to 0$, was (to two significant figures) equal to $-0.56$, precisely the value predicted independently by Lyne. We did not take our numerical results to a high enough value of $R_s$ to achieve this limit, although the leading-order vorticity profile along $\theta = 3\pi/2$ did tend to flatten out near $r = 1/2$ at the larger values of $R_s$. The large-$R_s$ solution to the perturbation problem (5.13)-(5.14) has not been obtained.

6. Physiological application

In the context of blood flow in the coronary arteries, relevant parameter values are (to one significant figure)

- Kinematic viscosity of blood $\nu = 4 \times 10^{-6}$ m$^2$ s$^{-1}$,
- Period of the heart beat $T \approx 1$ s.

We consider those extramyocardial arteries whose curvature is single-signed and varies relatively slowly with longitudinal position. For such arteries we may take (Guyton 1986)

- Radius $a = 2 \times 10^{-3}$ m,
- Radius of curvature $\hat{R} = 3 \times 10^{-2}$ m,
- Peak time–averaged blood velocity $U_p = 3 \times 10^{-3}$ m s$^{-1}$,
- Velocity amplitude $\hat{W}_p = 2 \times 10^{-1}$ m s$^{-1}$;

moreover, if we assume that the volume of the left ventricle is halved during systole, then the radius of curvature will be reduced by a factor $2^{-1/3} \approx 8 \times 10^{-1}$. From this, we find

$$\delta_0 = a/\hat{R} \approx 0.07, \quad \epsilon \approx 0.1.$$  \hfill (6.1)

Also

$$Re = \frac{U_p a}{\nu} \approx 150, \quad \alpha^2 = \frac{a^2}{\nu T} \approx 1.$$  \hfill (6.2)

We considered a dimensionless oscillatory pressure gradient:

$$-p_s = \dot{x}^2 W_p \cos (t + \Phi),$$  \hfill (6.3)

where $W_p$ is the dimensionless velocity amplitude that would be driven by such a pressure gradient in a straight pipe containing an inviscid fluid. Here the velocities were non-dimensionalized with respect to $v/a$ and so

$$W_p = a\hat{W}_p/v \approx 100.$$  \hfill (6.4)
350 S. L. Waters and T. J. Pedley

The secondary streaming Reynolds numbers, \( R_s \) and \( R_s^1 \), are defined as

\[
2R_s = \frac{(2\delta_0) W_p^2}{\alpha^2} = \frac{(2\delta_0)}{\alpha^2} \left( \frac{a\bar{W}_p}{v} \right)^2,
\]

\[
R_s^1 = (2\delta_0) \epsilon W_p = (2\delta_0) \epsilon \left( \frac{a\bar{W}_p}{v} \right)
\]

(cf. (2.32)) and so

\[
2R_s \approx 1320, \quad (2R_s)^{1/2} \approx 36, \quad R_s^1 \approx 1.4.
\]

The parameter \( R_t \), (2.33), is

\[
R_t = \alpha \epsilon \approx 0.1.
\]

When considering the flow driven by an oscillatory pressure gradient, the curvature was assumed to be small (\( \delta_0 \ll 1 \)) and to vary by a relatively small amount (\( \epsilon \ll 1 \)), assumptions that are reasonably well satisfied in the coronary arteries.

Initially \( x \) and \( R_s \) were kept order one compared with \( \delta_0 \) and \( R_t \). A small-parameter expansion was carried out in powers of \( \delta_0 \) followed by a small-parameter expansion in powers of \( R_s \), so that we assumed that \( \delta_0 \ll R_s \). The equations were then solved for asymptotically large and small values of the frequency parameter, \( x \).

From (6.2), (6.6) and (6.7), the assumption that \( x \) and \( R_s \) are order one compared with \( \delta_0 \) and \( R_t \) is reasonably well satisfied. Moreover, although \( \delta_0 \) is not much smaller than \( R_s \), it is smaller. The assumptions that \( x \ll R_t \) (low-frequency case) and \( R_t \ll 1/x^2 \) (high-frequency case) are approximately satisfied; however \( x \approx 1 \) and so is neither large nor small. In §5 results were found for small and intermediate values of \( R_s \). From (6.6) we see that the most physiologically significant solutions correspond to those found numerically for intermediate values of \( R_s \).

It appears that at present, the small-\( x \) case is most applicable to flow in the coronary arteries. For the small-\( x \) case, the magnitude of the ratio between the first perturbation to the time-dependent axial wall shear stress, due to the introduction of time-dependent curvature, (4.13), and the leading-order time-dependent axial wall shear stress (4.3) is, with \( s = O(1) \), approximately \( 8 \times 10^{-4} \). The corresponding ratio between the azimuthal wall shear stress perturbation (4.14) and the leading-order azimuthal wall shear stress (4.6) is \( 1 \times 10^{-2} \). Thus the predicted effect, for this particular range of parameters, is rather small. Our results are more significant for the light they shed on the qualitative effect of the time-dependent curvature on the distribution of wall shear stress. It was found that the Coriolis force terms, introduced into the equations by the time-dependent curvature, led to the generation of a steady secondary streaming perturbation. For low frequency, this steady streaming was towards the outside of the bend when the phase difference, \( \Phi \), between the curvature and the axial pressure gradient was such that \( \cos \Phi > 0 \) (the opposite being true when \( \cos \Phi < 0 \)). This perturbation to the steady streaming resulted in a correction to the axial velocity profile, with corresponding perturbations to the axial wall shear stress.

During systole, the increased intramyocardial pressure resulting from the contraction of the thick-walled left ventricle hinders blood flow into the heart muscle (but does facilitate venous drainage by compressing the coronary sinus and other veins). When diastole begins, the rate of flow rapidly increases to reach a maximum level in early diastole, from which it gradually declines. Now, when the ventricle contracts, the rotation of the artery is negative, and so \( \cos t > 0 \). The phase, \( \Phi \), must be such that the driving oscillatory pressure gradient rises initially and then starts to fall, as
the flow of blood is hindered by the increased intramyocardial pressure. At the start of diastole, the rotation of the artery is positive, so that \( \cos t < 0 \), and the phase difference, \( \Phi \), must be such that the driving pressure gradient rises rapidly and then gradually decreases. For this to be true, it is necessary that \( \pi/2 < \Phi < \pi \), so that \( \cos \Phi < 0 \). Thus for small \( \alpha \), we see that there is an increase in wall shear stress on the inside wall, which may be protective against atherosclerosis, and a decrease on the outside wall. Since atherosclerosis is more often found on the inside of the curve, this suggests that the prevalence of atherosclerosis would be worse without movement of the artery.

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